



РОССИЙСКИЙ ФЕДЕРАЛЬНЫЙ ЯДЕРНЫЙ ЦЕНТР ВНИИЭФ

# The use of high power lasers for studying rheological properties of substances

V.G. Rogachev (RFNC – VNIIEF)

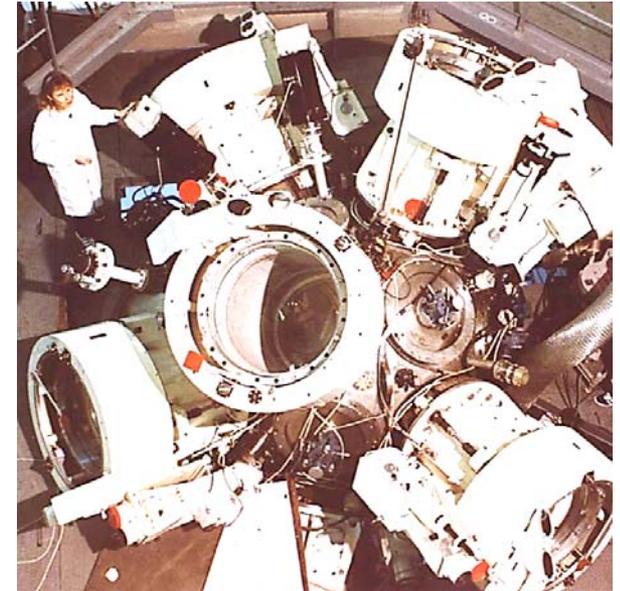
XIII Zababakhin's Scientific Hearings International Conference  
Snezhinsk, Russia, 20 - 24 March, 2017

- ✓ **Statement and results of laser experiments at RFNC-VNIIEF laser facilities on studying rheological properties of various substances are discussed.**
- ✓ **Main attention is devoted to experiments at low flows of laser radiation (LR) when relatively low intensity shock waves are generated within the target. The shock wave attenuation dynamics and spalls generation are investigated.**
- ✓ **The laser method simulating preventive nuclear explosion destructive effect on dangerous asteroids is considered.**
- ✓ **In laser experiments with stony asteroid micro models the criteria for chondrites asteroids effective destruction at contact nuclear explosion were defined.**

# "ISKRA-5" laser facility



**"ISKRA-5" building appearance**



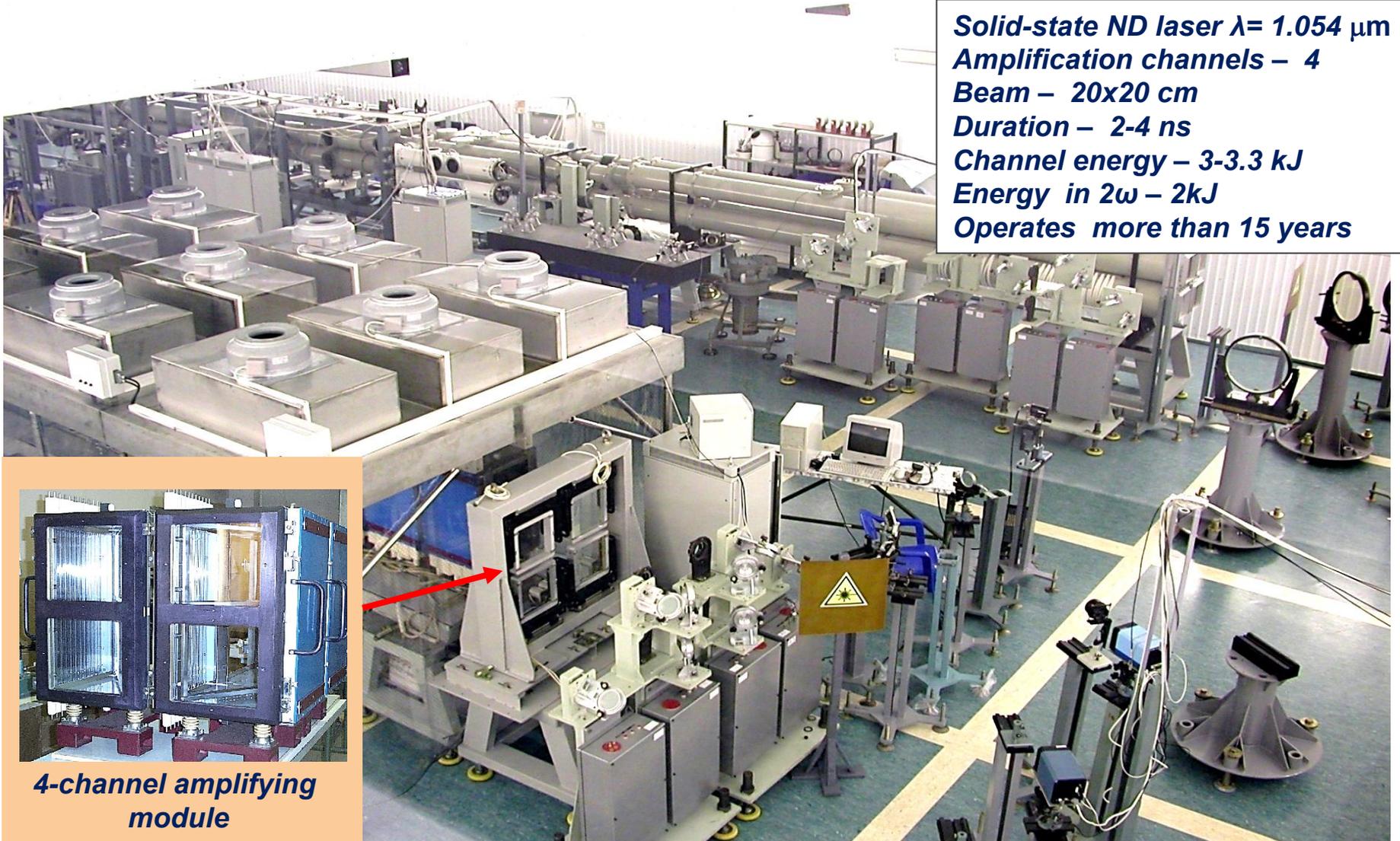
**12-beam spherical chamber  
Ø 2 m**

**12 – beam photo-dissociative iodine laser with  $\lambda = 1.315 \mu\text{m}$ ;  
Laser energy at first harmonics is up to 30 kJ  
Pulse duration – 0.4 ns  
Power – up to 100 TW**

# “LUCH” laser facility



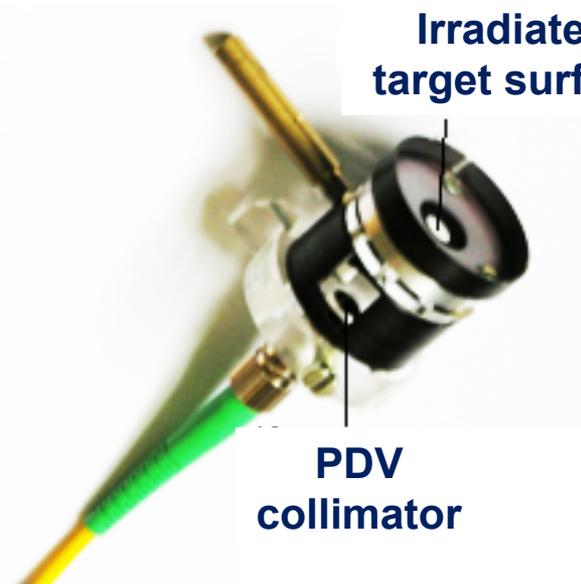
*Solid-state ND laser  $\lambda = 1.054 \mu\text{m}$   
Amplification channels – 4  
Beam – 20x20 cm  
Duration – 2-4 ns  
Channel energy – 3-3.3 kJ  
Energy in  $2\omega$  – 2kJ  
Operates more than 15 years*



**4-channel amplifying module**

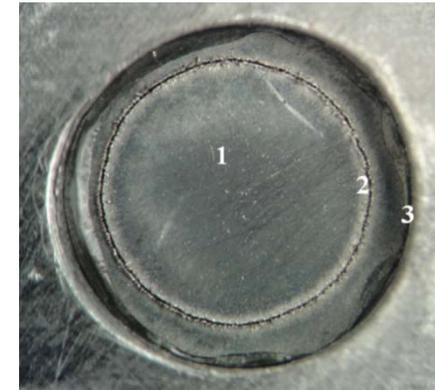
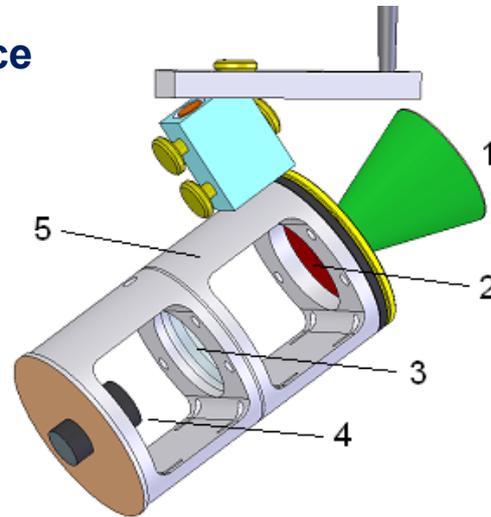
- Equations of state for various substances at pressures up to 70 MBar
- Soft XR mean free paths at temperatures up to 50 eV
- Soft X-radiation transport
- Gas dynamics instabilities and turbulent mixing
- Laser radiation transport within closed cavities
- LR absorption and reflection at flows  $10^{11} - 10^{14}$  W/cm<sup>2</sup>
- Methods of increasing laser radiation symmetry
- Magnetized plasma physics
- Physics of laser fusion targets of direct exposure
- Physics of laser fusion X-ray targets
- Non-equilibrium processes within inverted corona laser targets
- **Processes at moderate laser radiation flows**

# Rheological properties study of various substances at “LUCH” facility



Irradiated target surface

PDV collimator



## Target unit and draft photo

- 1 – high-power laser radiation beam;
- 2 – irradiated target backside;
- 3 – protective screen for collimator;
- 4 – PDV collimator;
- 5 – target unit case

## Photo of target backside surface initial state

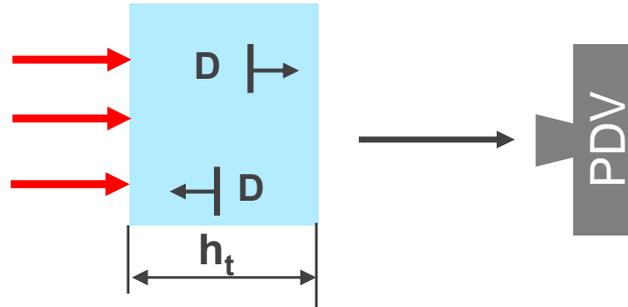
- 1 – polished target surface;
- 2 – circular groove, reducing edge effect at spallation;
- 3 – border of aperture within the case

**Photon Doppler Velocimetry (PDV) – laser optical - heterodyne method for velocity detecting.**

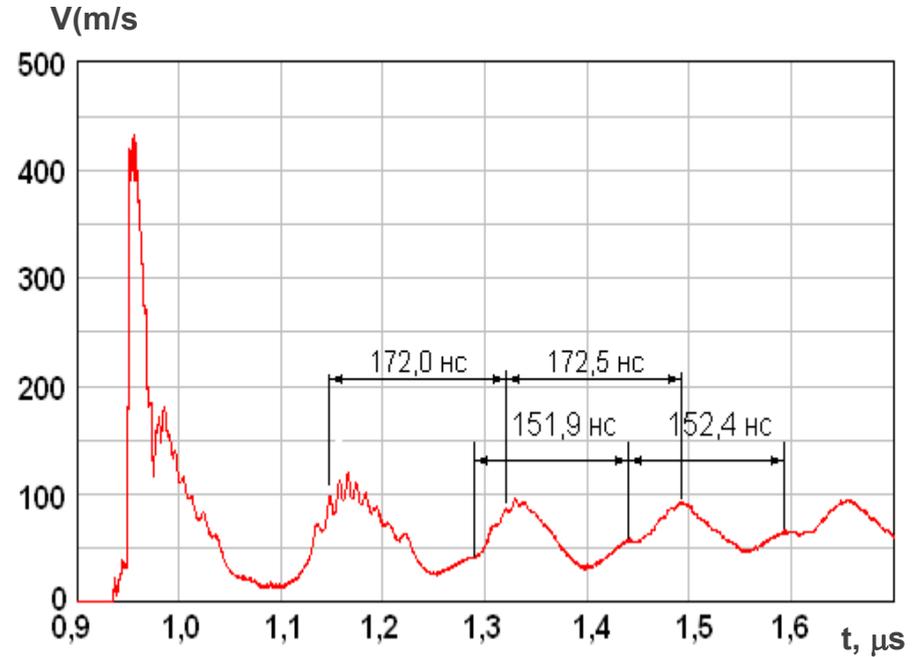
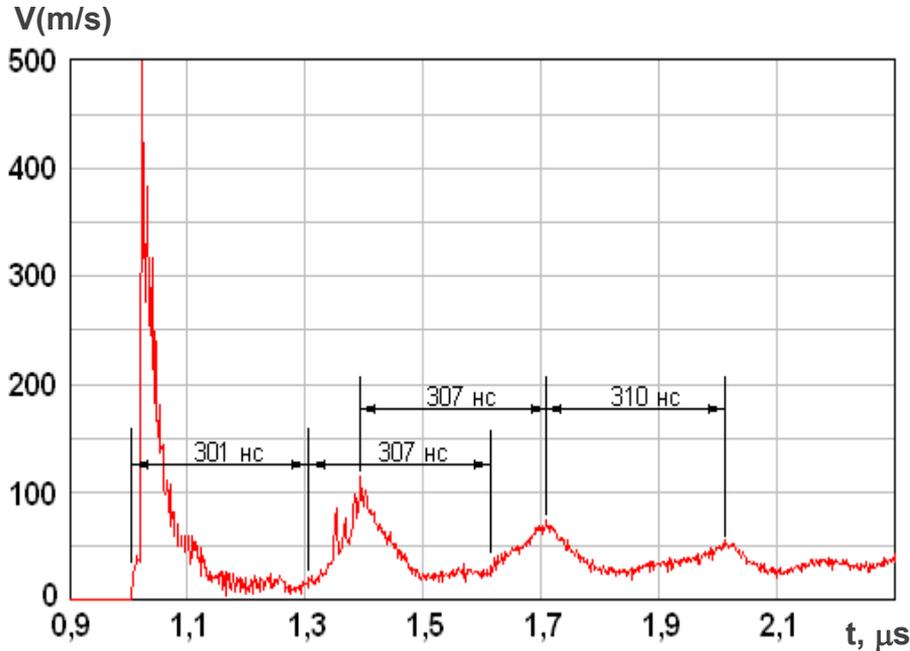
# Acoustic waves in alloy AL1060 (no spalls)



$h_t = 975 \pm 1 \mu\text{m}$   
 $Q_{LR} = 107 \pm 16 \text{ J}$   
 $\tau_{LR} = 5.0 \pm 0.3 \text{ ns}$   
 $S_{LR} = 9.5 \pm 0.9 \text{ mm}^2$   
 $\tau = 308.5 \pm 2 \text{ ns}$   
 $D = 6.32 \text{ km/s}$



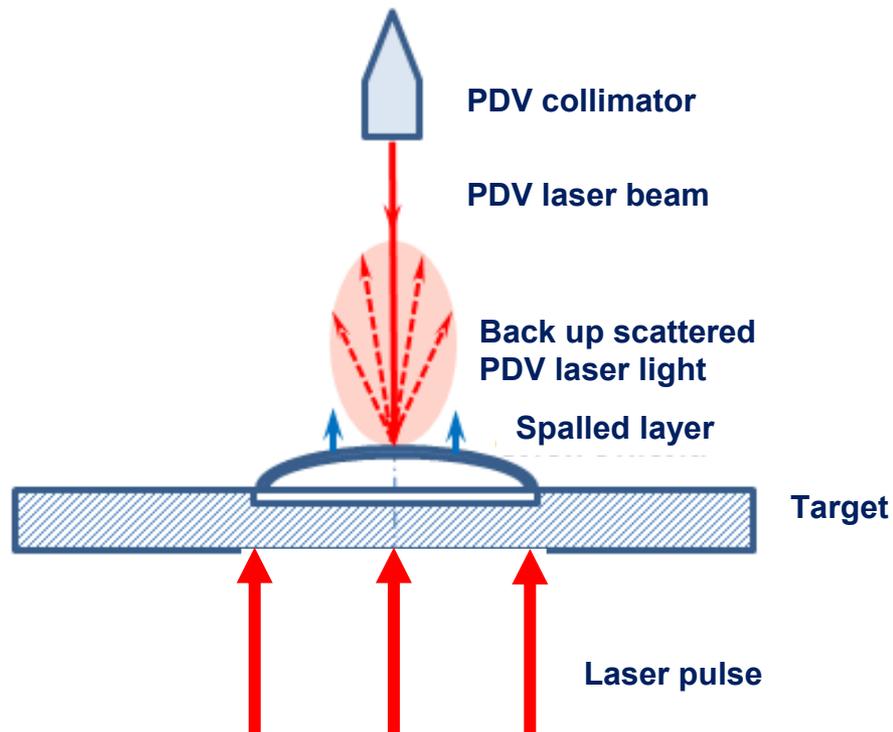
$h_t = 515 \pm 1 \mu\text{m}$   
 $Q_{LR} = 86 \pm 13 \text{ J}$   
 $\tau_{LR} = 4.5 \pm 0.1 \text{ ns}$   
 $S_{LR} = 9.5 \pm 0.9 \text{ mm}^2$   
 $\tau = 172.2 \pm 1 \text{ ns}$   
 $D = 5.98 \text{ km/s}$



**Velocity of AL1060 flat target back surface vs time.  
 Attenuation of low shock (acoustic) waves.**

# Spalls in AL1060 alloy

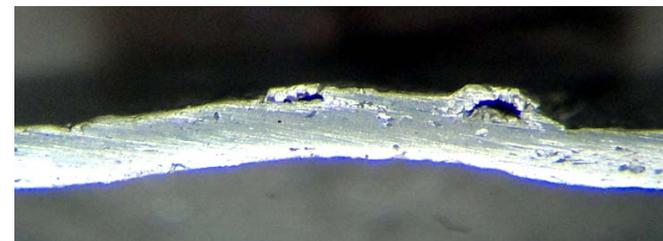
(Experiments at “LUCH” facility)



## Target cross-section photos



no spalls



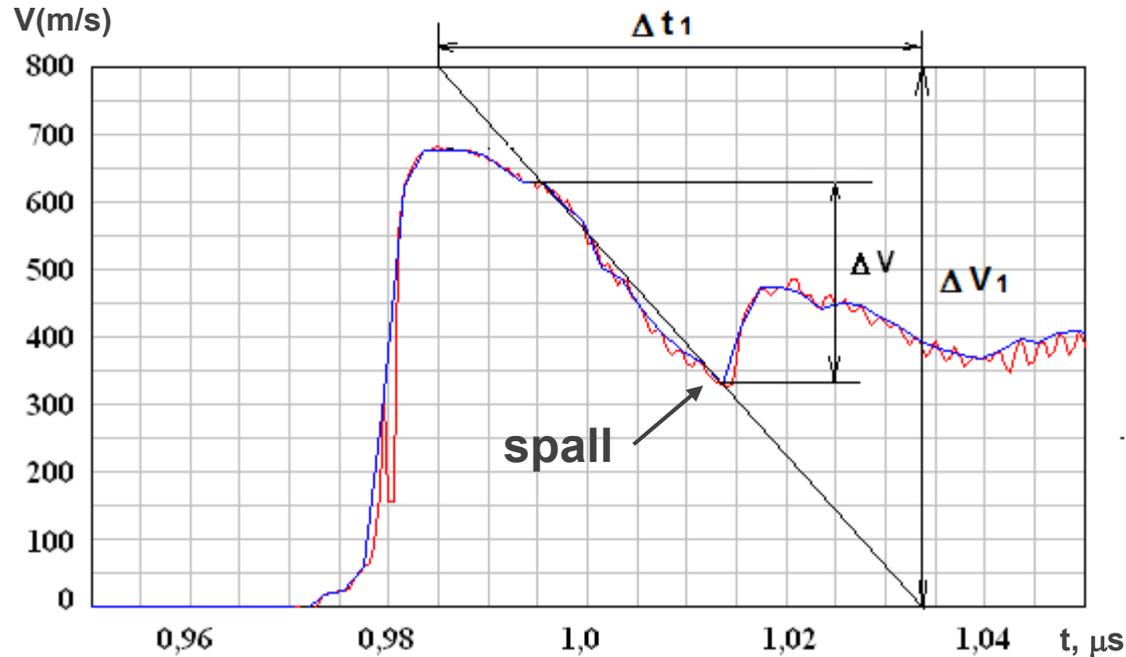
Concealed spalls



Complete spall

“LUCH” facility experimental layout  
(flat liner)

# Target back surface velocimetry (spalls)



**Influence of averaging interval in Hilbert transform on results of PDV  
oscillogram numerical processing :**

$\Delta\tau = 125$  ps – red;  $\Delta\tau = 2000$  ps – blue;  $d\varepsilon/dt = -0.5 \cdot dv/dt/D$ .

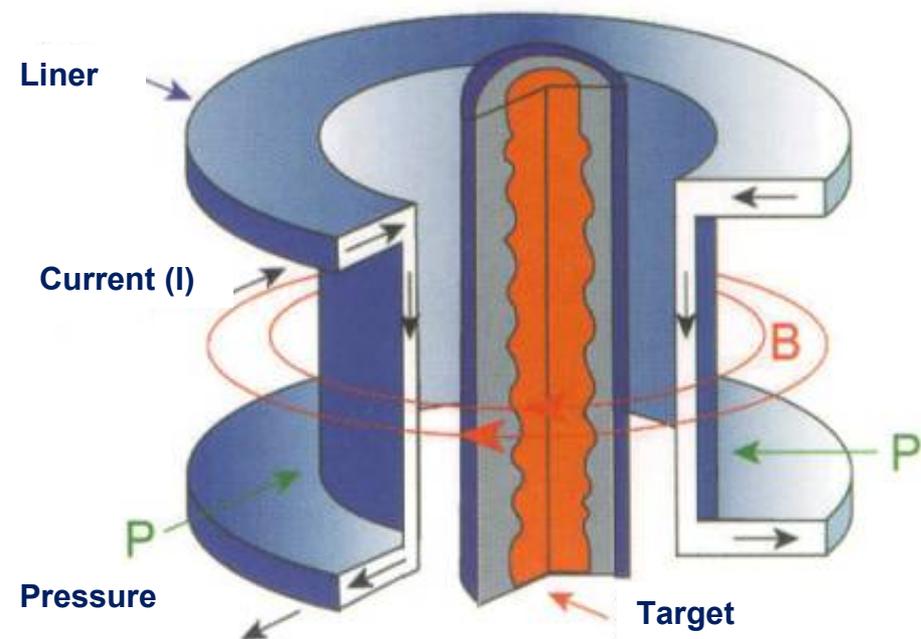
# Spalls in AL1060 alloy

(electro physical experiments with magneto-cumulative generator)



## Cylindrical liner

### Loading scheme



Spall origin



Complete spall

ДОКЛАДЫ АКАДЕМИИ НАУК, 2013, том 448, № 3, с. 285–288

ФИЗИКА

УДК 539.4

## ИССЛЕДОВАНИЕ РЕОЛОГИЧЕСКИХ СВОЙСТВ АЛЮМИНИЯ С ПРИМЕНЕНИЕМ ВЗРЫВОМАГНИТНЫХ ГЕНЕРАТОРОВ

© 2013 г. В. А. Васюков, А. М. Глыбин, П. В. Дудай, В. И. Дудин, А. А. Зименков, В. А. Иванов, А. В. Ивановский, А. И. Краев, А. И. Кузьев, С. С. Надежин, А. А. Петрухин, А. Н. Скобелев, О. А. Тюпанова, W. L. Atchison, D. B. Holtkamp, A. M. Kaul, R. E. Reinovsky, G. Rodrigues, L. J. Tabaka, C. L. Rouscalp, J. B. Stone, D. M. Oro, M. Salazar, J. R. Griego, J. R. Payton, D. T. Westley

Представлено академиком Р.И. Ильяевым 30.03.2012 г.

Поступило 18.07.2012 г.

# Aluminum alloys strength properties (comparison of the results)



**Purpose:** extension of experimental data range on various material properties

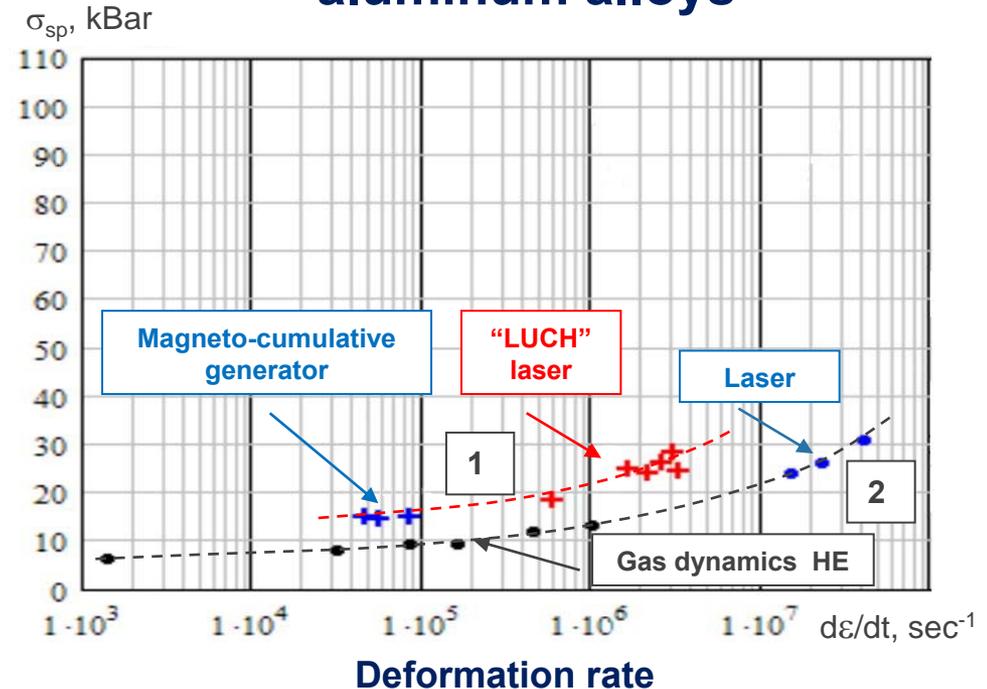
**Methods:** laser pulse and electro physical loading of liners, made from materials under consideration

**Particular features:** mutual verification of the methods, scale factor, enhanced range of effective parameters

**Current stage:** working-out of methods and their optimization

**Expected results:** creating of experimental material properties database for pulse processing techniques

## Spall strength measurement for aluminum alloys

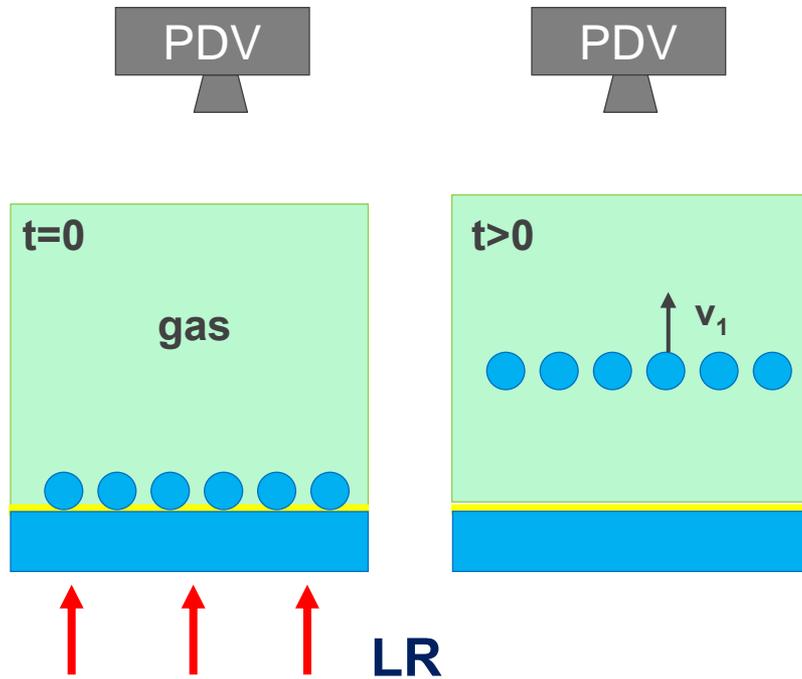


1 – nano-dispersed aluminum AL1060.

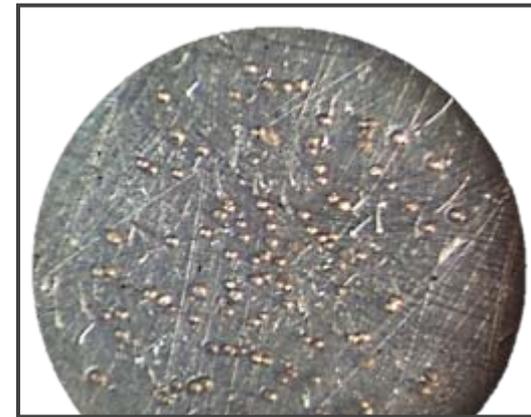
2 – aluminum alloy AMr6M –

A.I.Bushman, G.I.Kannel, A.L.Ni, V.E.Fortov “  
Thermal physics and dynamics of intensive pulse  
actions”, Chernogolovka, 1988.

# PDV technique calibrating at “LUCH” facility experiments



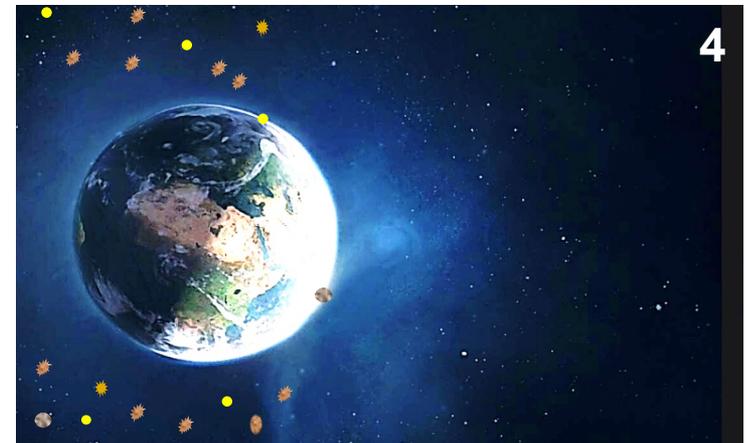
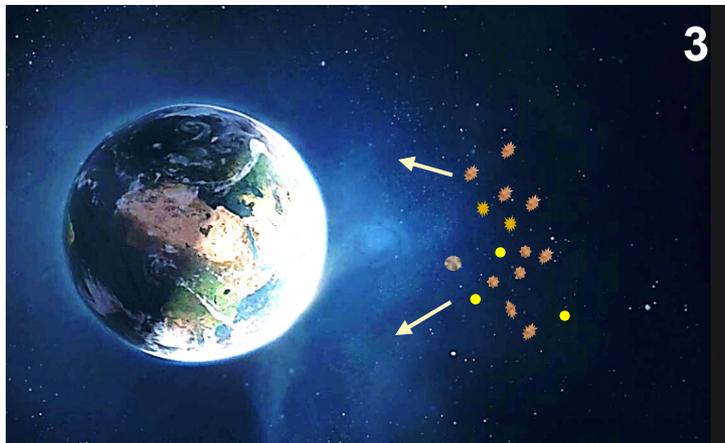
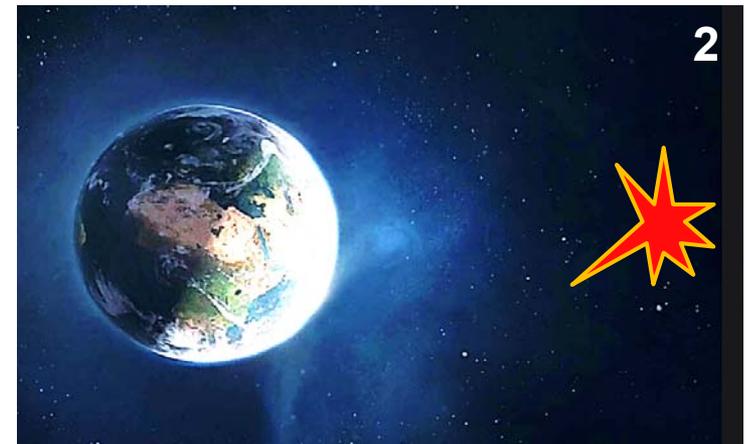
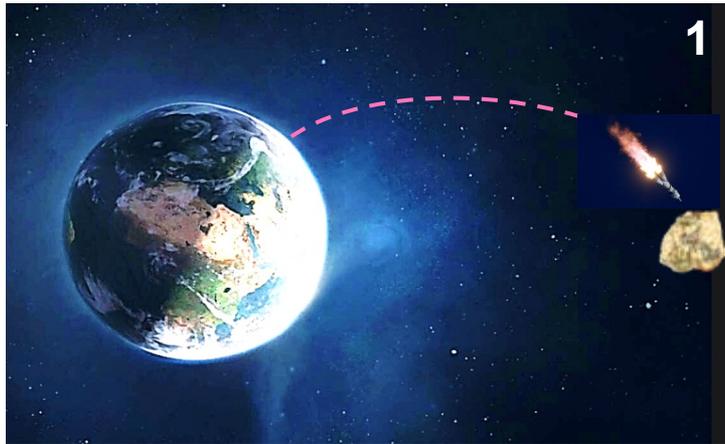
Experimental layout



Lead target with small spheres (experimental sample)

Small spheres deceleration in gas ( $\varnothing \sim 100\mu\text{m}$ );  $\Delta p < \sigma_{sp}$

# Asteroid safety



**Asteroid is disintegrated under nuclear explosion with small fragments, scattering wide apart . Catastrophe is prevented.**

# Asteroid safety

(on navigation)

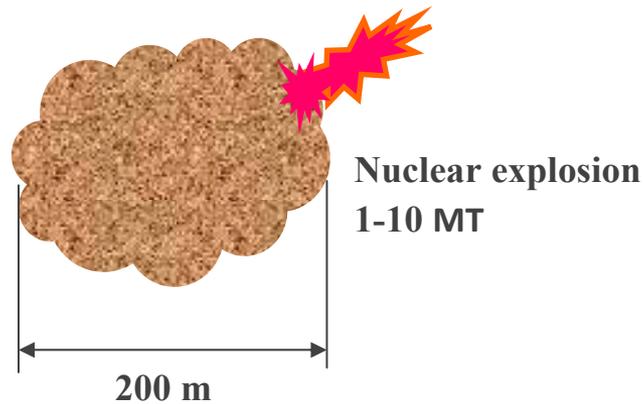


## Estimation of nuclear protection conditions:

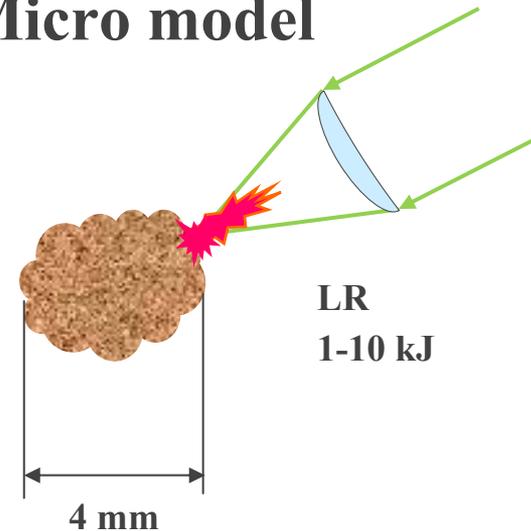
- When velocity of asteroid fragments at its disintegration is  $V_{fr} \sim 10$  m/s and Asteroid - Earth typical approaching velocity is  $V_{ast} \sim 20$  km/s, it is necessary to carry out nuclear explosion at distance  $L = RV_{ast}/V_{fr} \sim 13$  Mln. km to deviate fragments on safe distance, which is more than Earth radius  $R_{Earth} = 6400$  km.
- For asteroids with 200 m in diameter self gravity of fragments is negligible, if its velocities  $V_{fr}$  are greater than 1 m/s.
- The rocket must start  $t = L/V_r \sim 15$  days before asteroid arrival at point of explosion with velocity  $V_r \sim 10$  km/s.
- At that moment the asteroid will be at the distance  $L_1 = L(V_{ast} + V_r)/V_r \sim 39$  Mln. km from the Earth.
- For more accuracy Kepler orbits must be used.
- High navigation accuracy should be provided during the whole flight period to ensure controlled effect of danger elimination.

- The method, simulating high-power nuclear explosion destructive effect on asteroids, is introduced and realized. It is based on physical and geometrical similarity concept.
- It is proposed to replace real asteroid of hundreds meters in diameter by its scale model, fabricated from the similar substance with the similar physical properties and having a few mm in diameter.
- Nuclear explosion energy release (in megatons) on the asteroid surface should be replaced by laser energy pulse release (hundreds J) on small area of the scale model surface.
- Hydrodynamic effects and processes of spallation will be physically similar both in full-scale and model scale tests, when explosion (laser pulse) energy to asteroid (model) mass ratios are approximately equal.
- On "ISKRA-5" and "LUCH" laser facilities the criteria for stony asteroid destruction at contact nuclear explosion were experimentally estimated using physical similarity principle.

## Asteroid



## Micro model



**Large stony asteroids (80-90% from total number) are breakable and have low destruction threshold (0.05-0.1 kBar). It lets us hope on their effective splitting with the help of nuclear explosion.**

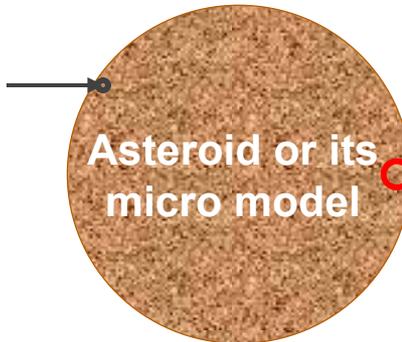
**Equality of effecting energy to object mass ratios is key criterion of the physical similarity of processes in both full-scale and model scale tests.**

# Asteroid safety

(2D – calculation similarity confirmation)

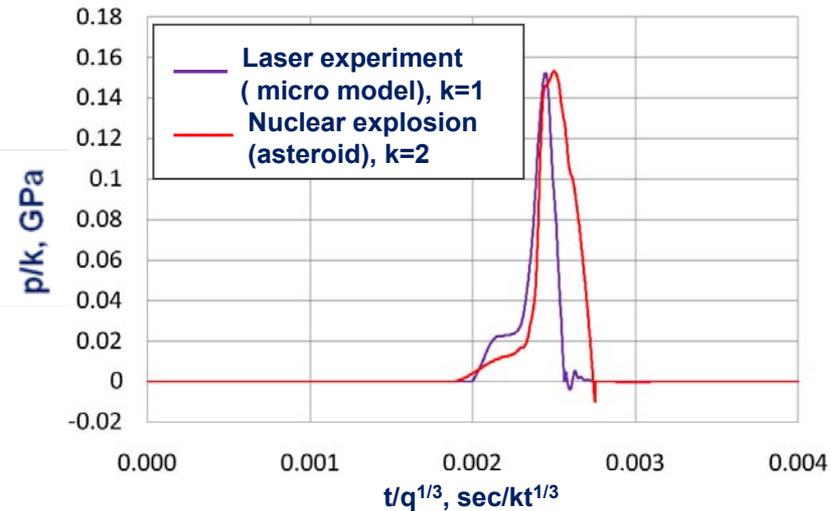
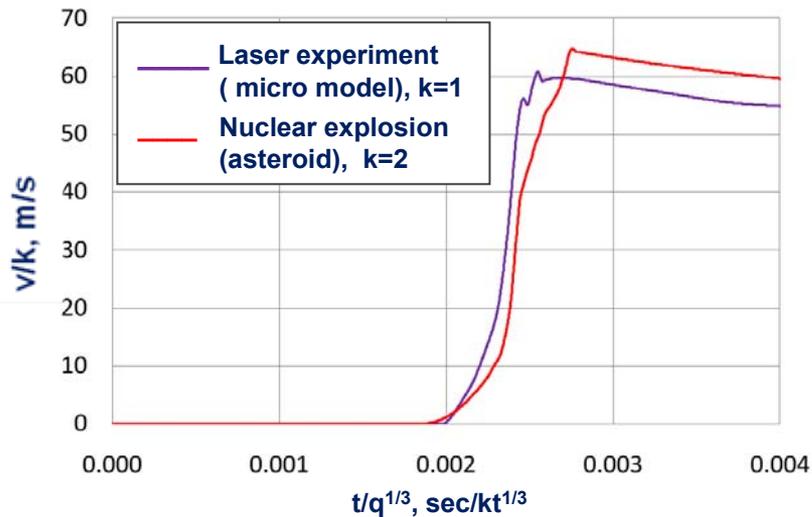


Observation point



Explosion

## Calculation layout

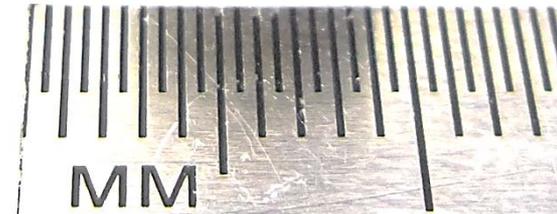


The example of mass velocity and pressure dependence on normalized time at the presented observation point (kt – kilotons).

# Asteroid safety



**Chelyabinsk asteroid fragment**  
**Mass ~ 0.5 tons**  
**(initial mass ~ 10000 tons)**



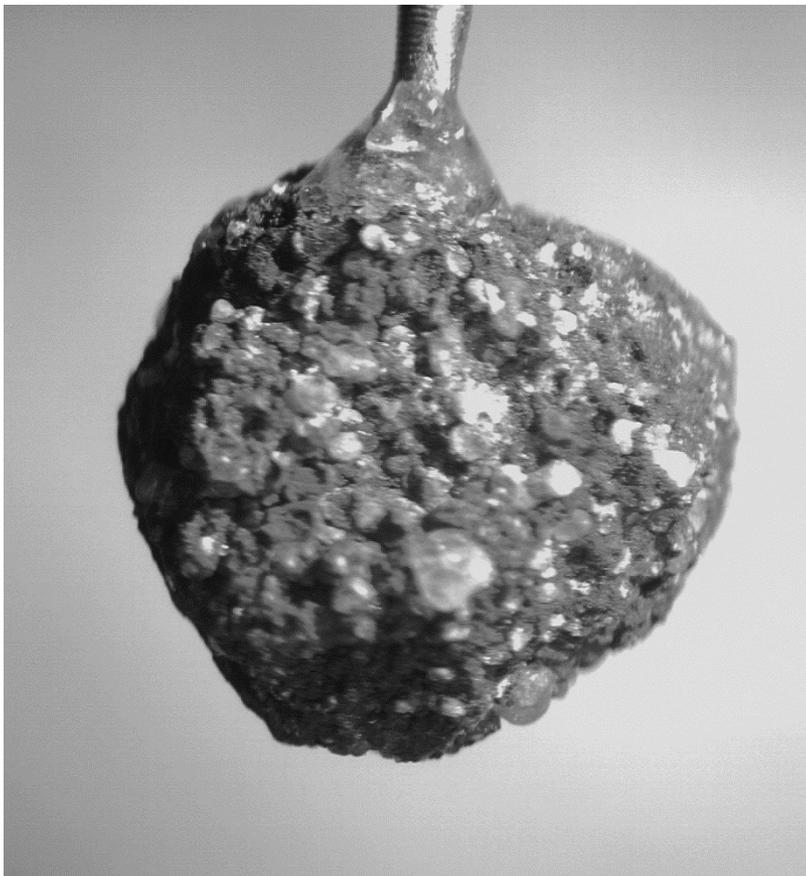
**Stony asteroid micro model,**  
**manufactured at VNIIEF,**  
**Mass ~ 1 gram.**

# Asteroid safety

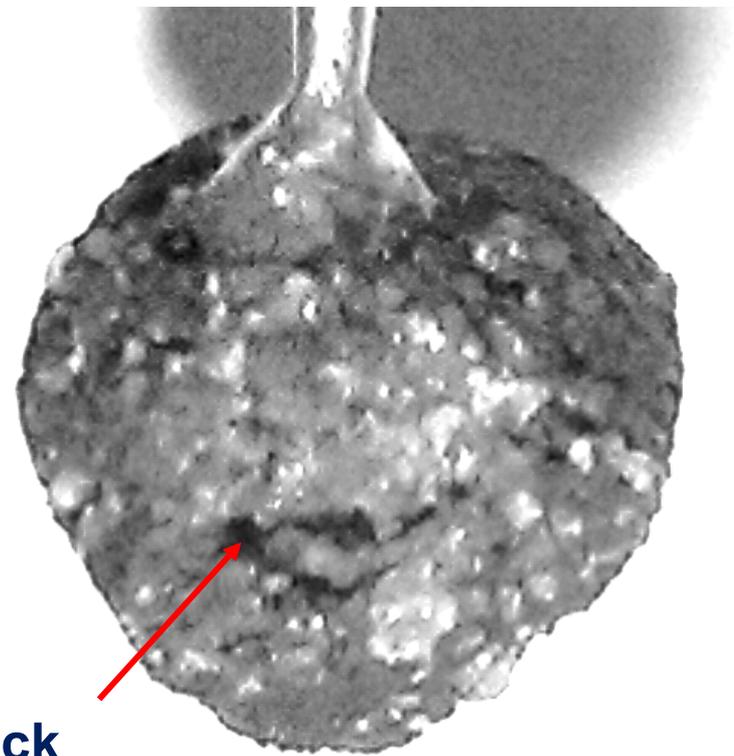
(laser simulation, "ISKRA-5")



Photos of target  $\varnothing$  8 mm at "ISKRA-5" laser experiment  
(at destruction threshold)



Before experiment

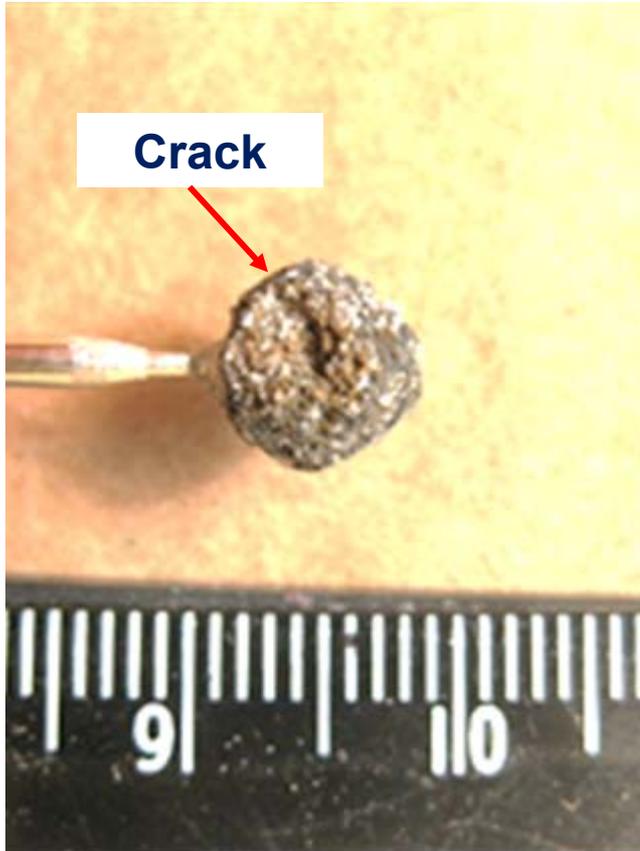


Crack

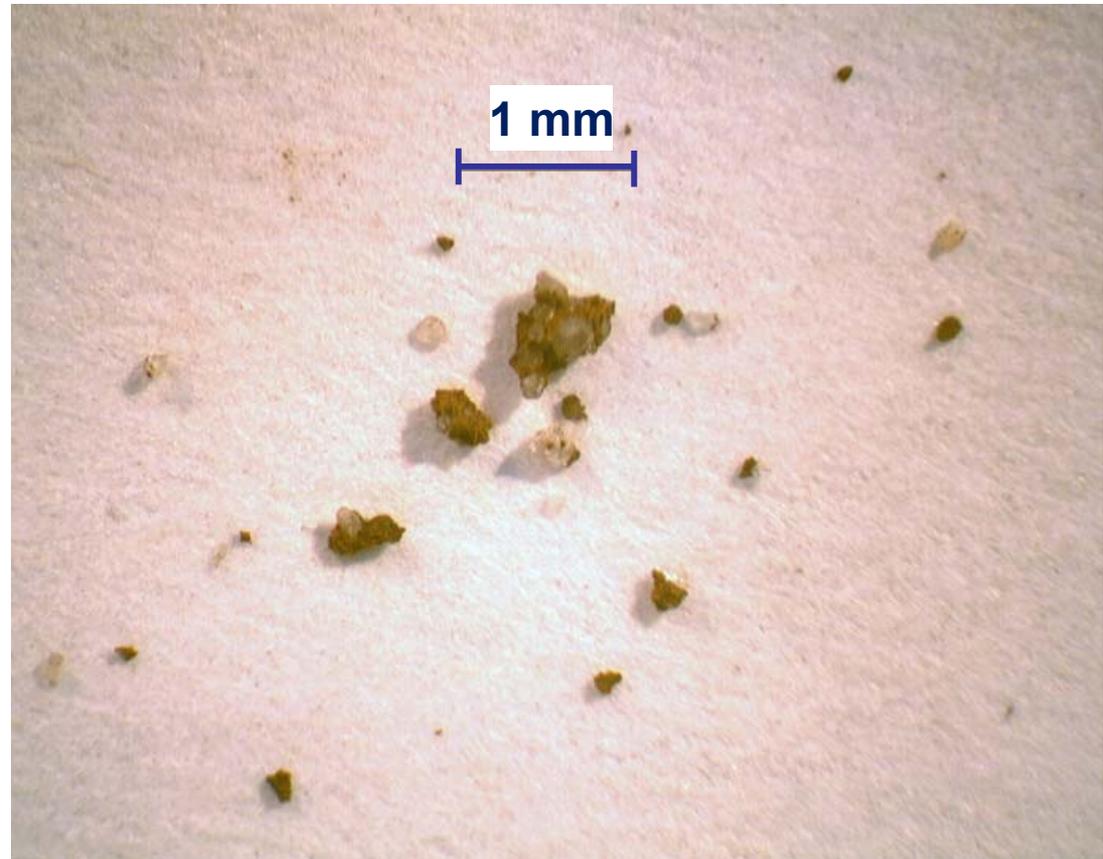
After experiment

# Asteroid safety

(laser simulation, "ISKRA-5")



Target  $\varnothing$  8 mm  
after experiment.



Fragments of target  $\varnothing$  4 mm,  
gathered after experiment.

# Asteroid safety

(on artificial asteroid manufacturing)



## 1. Artificial chondrite: chemical composition

$\text{SiO}_2$  – 40%  
 $\text{FeS}$  – 6%  
 $\text{MgO}$  – 26%  
 $\text{Fe}_2\text{O}_3$  – 18%

## 2. Initial materials splitting up ranges: 1-20-800 $\mu\text{m}$ .

## 3. A few sets of mini asteroids with different compression strength $\sigma$ are manufactured :

63 Bar

225 Bar

900 Bar

1341 Bar

## 4. Maximal tension strength is 10 times less, accordingly.

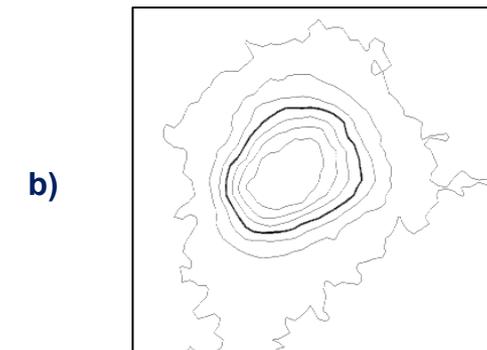
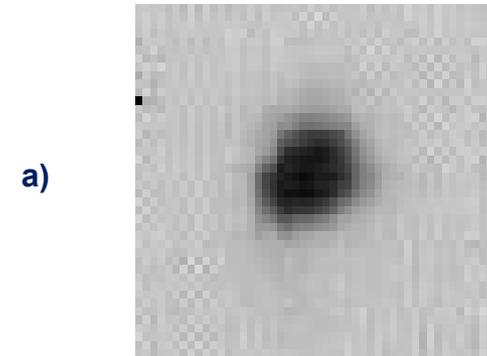
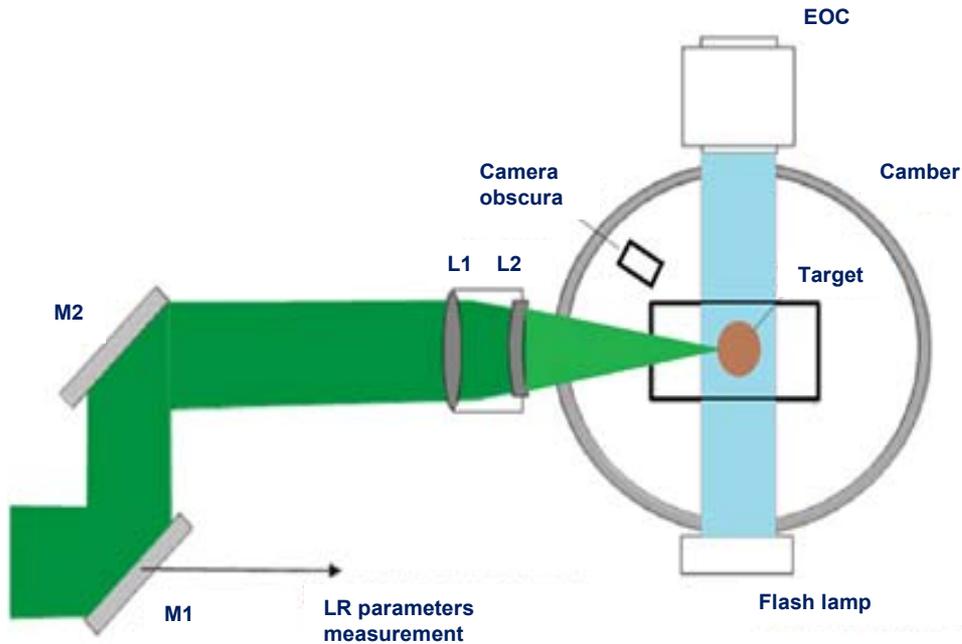


Typical diagram of artificial  
asteroid material sample  
compression

Maximal tension strength  
 $\sigma_{sp} = 6,3 \text{ MPa}$  (63 Bar), and  
deformation  $\varepsilon_{sp} = 5,5 \%$ .

# Asteroid safety

(Experiments at “LUCH” facility)



**Experimental set-up layout :**  
M1, M2 – mirrors, L1, L2 – lenses,  
EOC – Electron Optical Converter.  
 $\lambda=0.527 \mu\text{m}$ ,  $E_{\text{las}}= 100 \div 500 \text{ J}$ ,  
 $\tau_{0,5} = 1.4 \div 2.2 \text{ ns}$ .

**Typical “LUCH” experiment:**  
a) X-ray camera obscura photo  
of the laser spot.  
b) X - ray intensity distribution isolines.  
 $d_{\text{eff}}=129 \mu\text{m}$ .

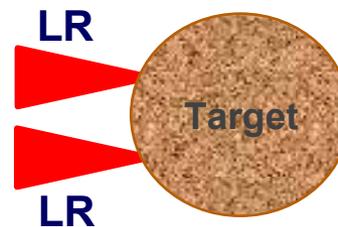
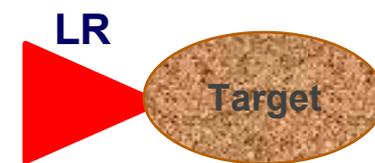
# Asteroid safety

(variation of parameters)



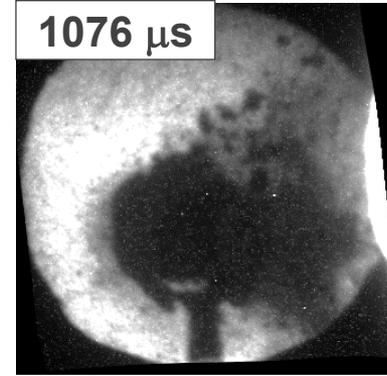
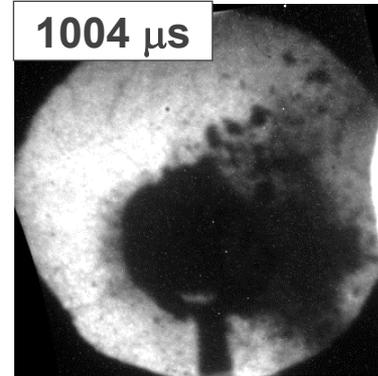
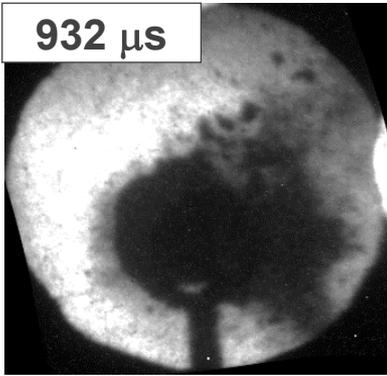
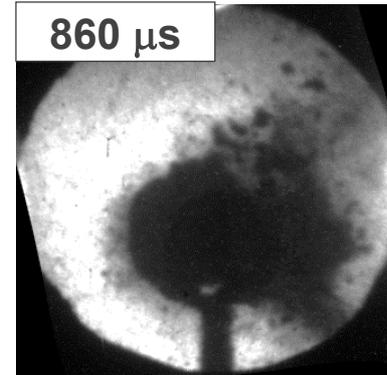
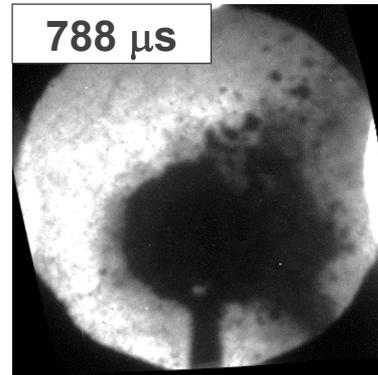
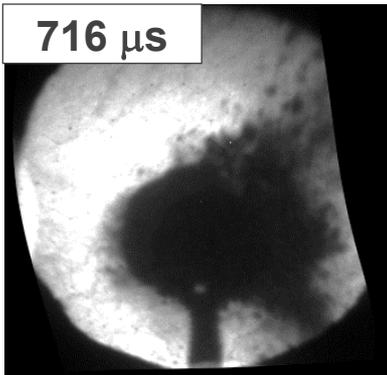
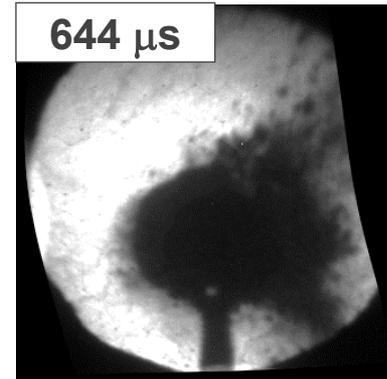
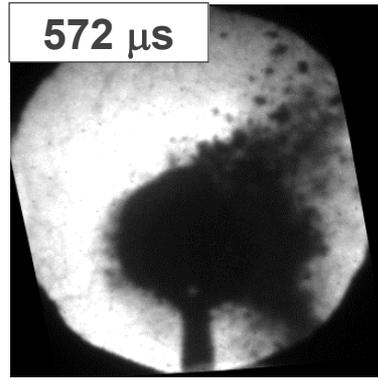
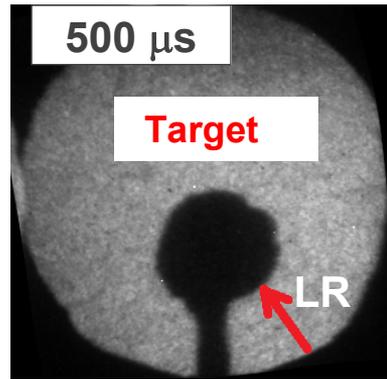
To determine the criteria for asteroids effective destruction at nuclear explosion the following laser experiments were carried-out :

- With variation of mini asteroids size and shape
- With variation of the substance heterogeneous structure
- With micro models having different compression strength
- With the natural chondrite asteroid substance
- With variation of energy release place
- With imitation of nuclear explosion in a cavity



# Asteroid safety

(target fragments dispersion at “LUCH” facility)



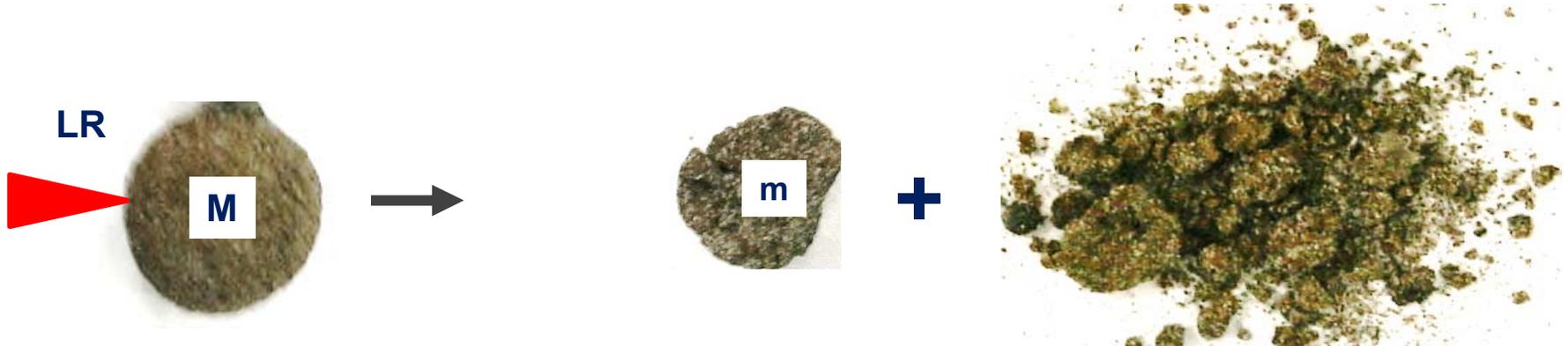
Typical dynamics of target fragments dispersion.  
Experiment 05.02.15.  
Start of shooting – 500  $\mu\text{s}$  after target irradiation.

# Asteroid safety

(“LUCH”, **sphere**,  $\sigma = 225$  Bar )



29.09.2015 E/M= 500 J/g. Partial destruction. Backside super large fragment  $m = 0.35 \cdot M$ .



23.09.2015 E/M= 750 J/g. No super large fragments.



# Asteroid safety

(“LUCH”, sphere,  $\sigma=1341$  Bar )



21.07.2015 E/M= 740 J/g. Partial destruction. Backside super large fragment m= 0.75·M.



22.07.2015 E/M= 1100 J/g. No super large fragments.



# Asteroid safety

(“LUCH”, **ellipsoid**  $\sigma=225$  Bar )



22.01.2016 E/M= 1300 J/g. Partial destruction. Backside super large fragment m= 0.45·M.



03.02.2016 E/M= 2900 J/g. Partial destruction. Backside super large fragment m= 0.23·M.

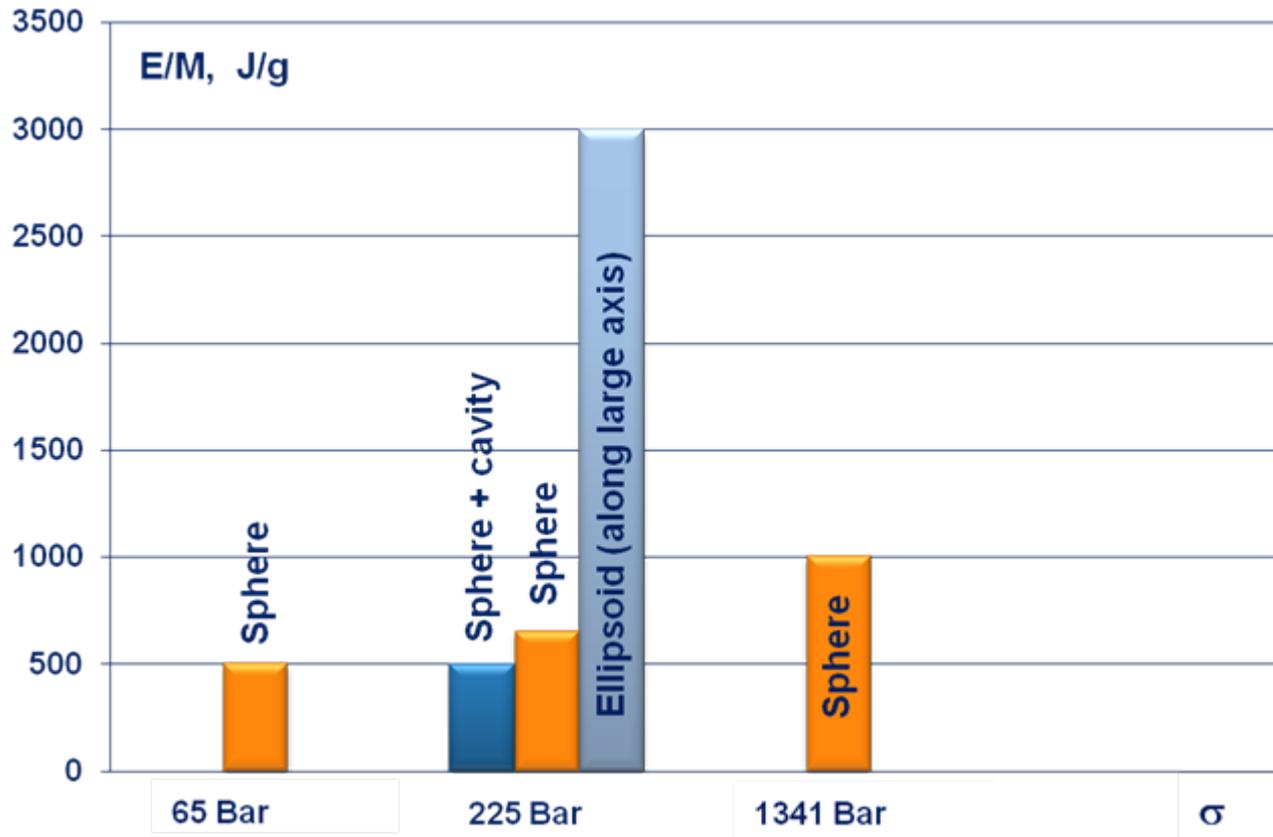


03.02.2016 E/M= 3000 J/g. No super large fragments.



# Asteroid safety

(“LUCH” facility experimental summary)



**Dependence between asteroid micro model destruction criterion and compression strength.**

- The direct numerical calculations confirmed quantitatively the existence of similarity.
- Stony asteroid effective destruction takes place at nuclear explosion energy to mass ratio more than  $E/M = 1000 \div 3000 \text{ J/g}$ .
- Laser experiments results converted into full-scale conditions demonstrate that nuclear explosion with energy  $\sim 6 \text{ Mt}$  destructs effectively 200 m diameter stony asteroid.
- Effective destruction threshold rises for ellipsoid-shape asteroids.
- Typical velocity of macroscopic fragments is the same both in full-scale and model cases. Fragments velocities were about  $V_{fr} \sim 4.5 \div 15 \text{ m/s}$ .



РОСАТОМ

РОССИЙСКИЙ ФЕДЕРАЛЬНЫЙ ЯДЕРНЫЙ ЦЕНТР ВНИИЭФ

**Thank you for your kind  
attention!**